

Chapter 7. Quantum Theory and Atomic Structure

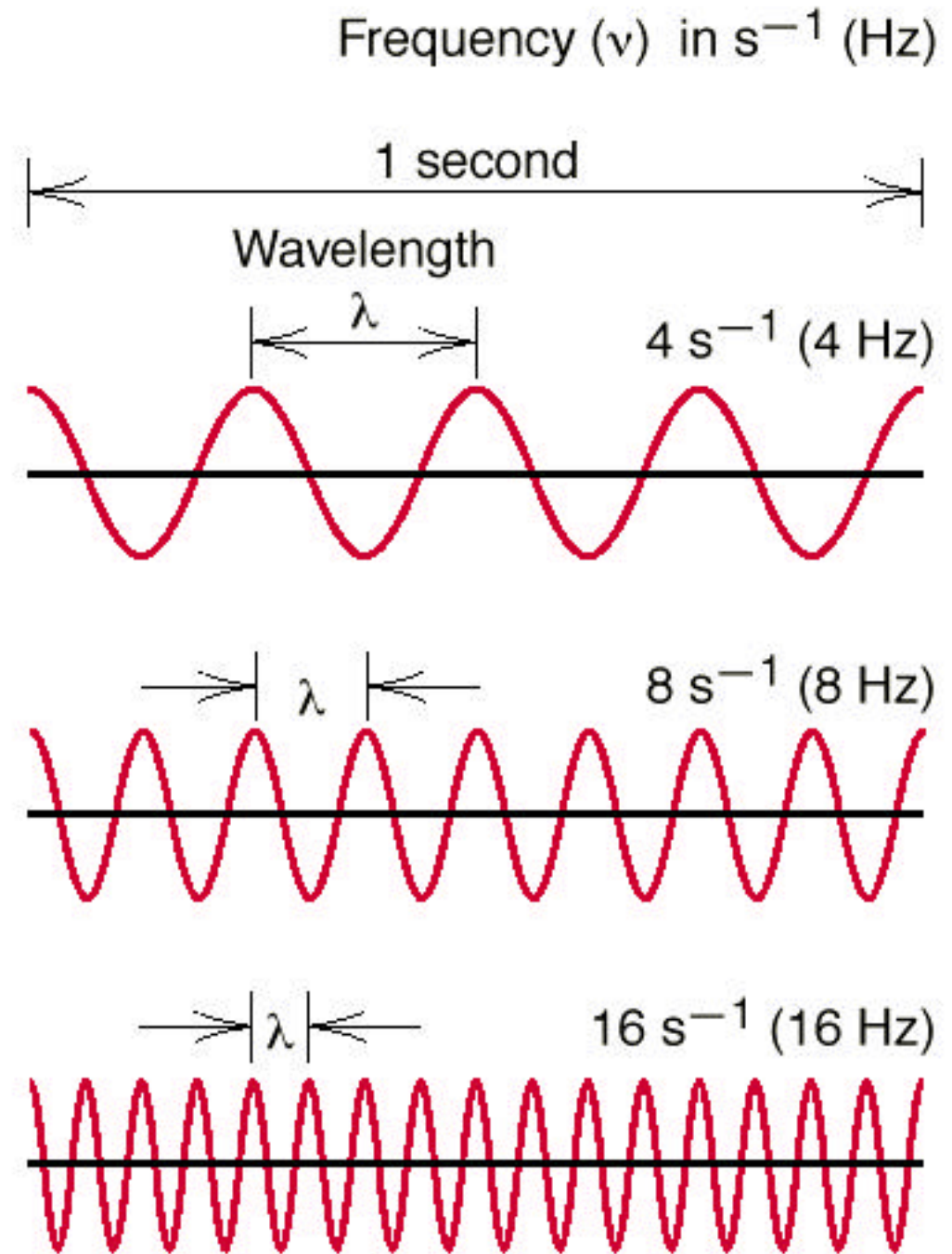
- A problem arose in Rutherford's nuclear model. A nucleus and electron attract each other; to remain apart the electron must move.
- The energy of the electron's movement must balance the energy of attraction.
- Physics established that a charged particle moving in a **curved path** such as an electron in an atom must give off energy.
- Why doesn't the electron continuously lose energy and spiral into the nucleus?
- Subatomic matter seemed to **violate common experience**. A new model was needed to describe atoms.

Concept 7-1. The wave characteristics of light (frequency, wavelength, and speed; the meaning of amplitude) and a general overview of the electromagnetic spectrum

- Visible light is one type of electromagnetic (EM) radiation.
- Electromagnetic radiation travels as waves—the result of oscillating electric and magnetic fields moving simultaneously through space.
- The wave properties of electromagnetic radiation are described by two interdependent variables, frequency and wavelength.



- Frequency (ν , Greek nu) is the number of cycles the wave makes per second, expressed in units of 1/second (s^{-1}), or hertz (Hz).



- Wavelength (λ , Greek lambda) is the distance between any point on a wave and the corresponding point on the next wave.

is expressed in meters, nanometers (nm, 10^{-9} m), picometers (pm, 10^{-12} m), and angstroms (\AA , 10^{-10} m).

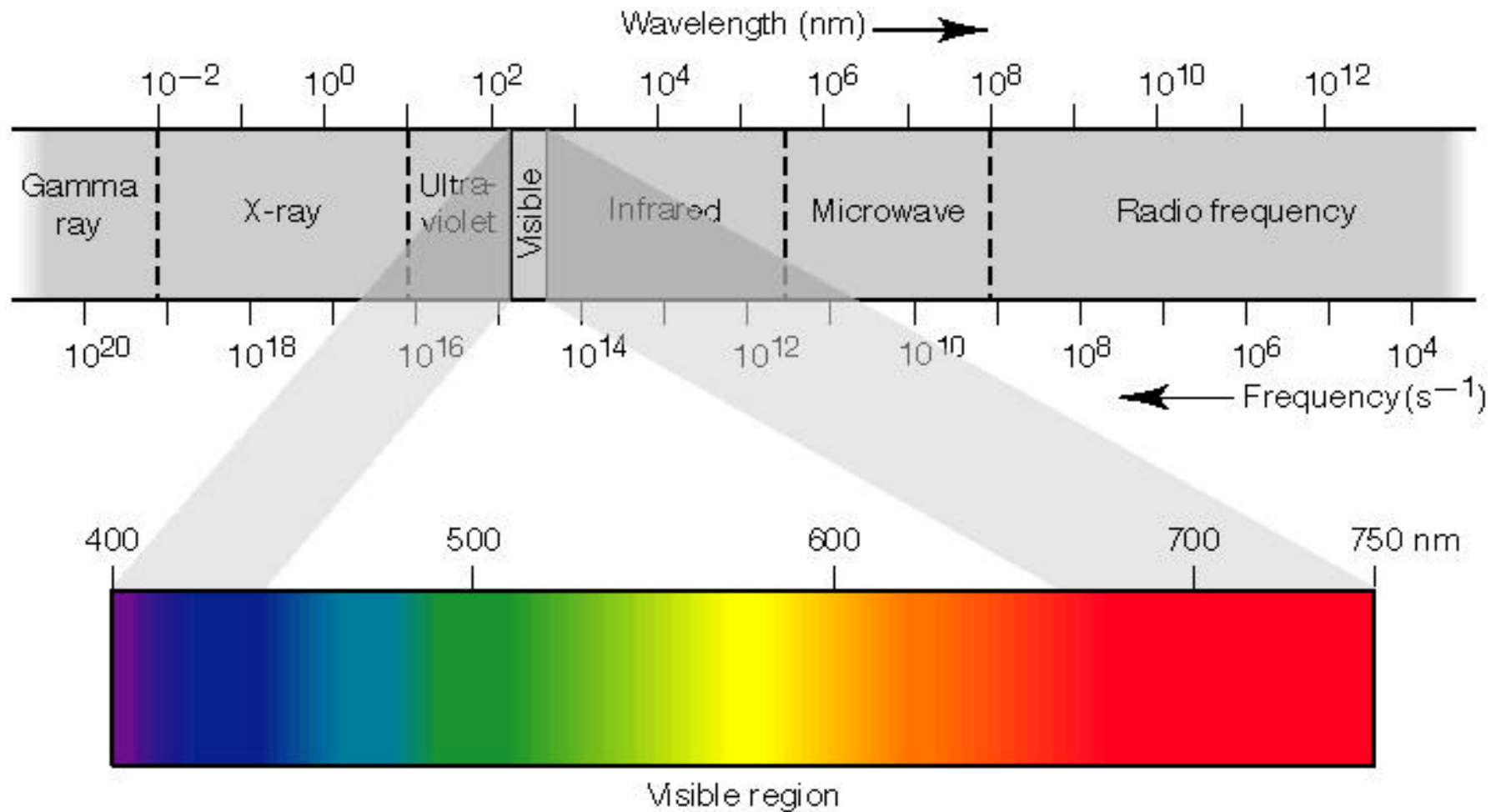
- The **speed** of the wave (m/s) is the product of its frequency (cycles per second) and its wavelength (meters per cycle):

$$\text{Speed} = \frac{\text{cycles}}{\text{s}} \times \frac{\text{m}}{\text{cycle}} = \frac{\text{m}}{\text{s}}$$

- In a vacuum, all electromagnetic radiation travels at **3.00×10^8 m/s**, a constant called the speed of light (c):

$$c = \nu\lambda$$

- Radiation of high frequency has a short wavelength, and vice versa.



Skill 7-1. Interconverting frequency and wavelength

Problem: A dentist uses x-rays ($\lambda = 1.00 \text{ \AA}$) while a patient listens to an FM radio station ($\lambda = 325 \text{ cm}$) and looks out the window at blue sky ($\lambda = 473 \text{ nm}$). What is the frequency (s^{-1}) of the EM from each?

Plan:

$$\begin{aligned}1 \text{ \AA} &= 10^{-10} \text{ m} \\1 \text{ cm} &= 10^{-2} \text{ m} \\1 \text{ nm} &= 10^{-9} \text{ m}\end{aligned}$$

$$\nu = \frac{c}{\lambda}$$

Wavelength (given units)



Wavelength (m)



Frequency (s⁻¹)

Solution: For x-rays. Converting from Angstroms to meters:

$$\lambda = 1.00 \text{ \AA} \times \frac{10^{-10} \text{ m}}{1 \text{ \AA}} = 1.00 \times 10^{-10} \text{ m}$$

Calculating the frequency:

$$\nu = \frac{c}{\lambda} = \frac{3.00 \times 10^8 \text{ m/s}}{1.00 \times 10^{-10} \text{ m}} = 3.00 \times 10^{18} \text{ s}^{-1}$$

For the radio station. Combining steps to calculate the frequency:

$$\nu = \frac{c}{\lambda} = \frac{3.00 \times 10^8 \text{ m/s}}{325 \text{ cm} \times 10^{-2} \text{ m/1 cm}} = 9.23 \times 10^7 \text{ s}^{-1}$$

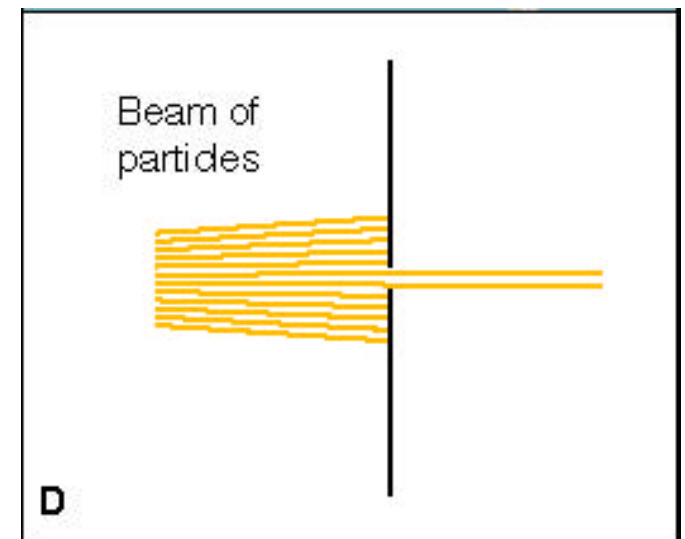
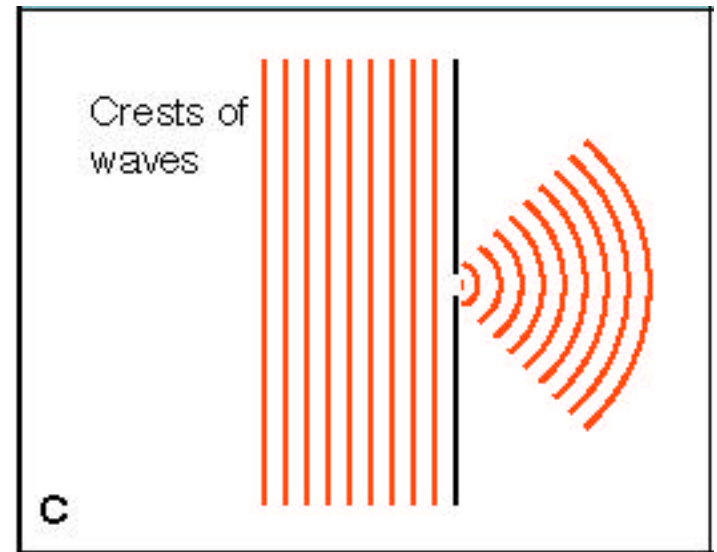
For the sky. Combining steps to calculate the frequency:

$$\nu = \frac{c}{\lambda} = \frac{3.00 \times 10^8 \text{ m/s}}{473 \text{ nm} \times 10^{-9} \text{ m/1 nm}} = 6.34 \times 10^{14} \text{ s}^{-1}$$

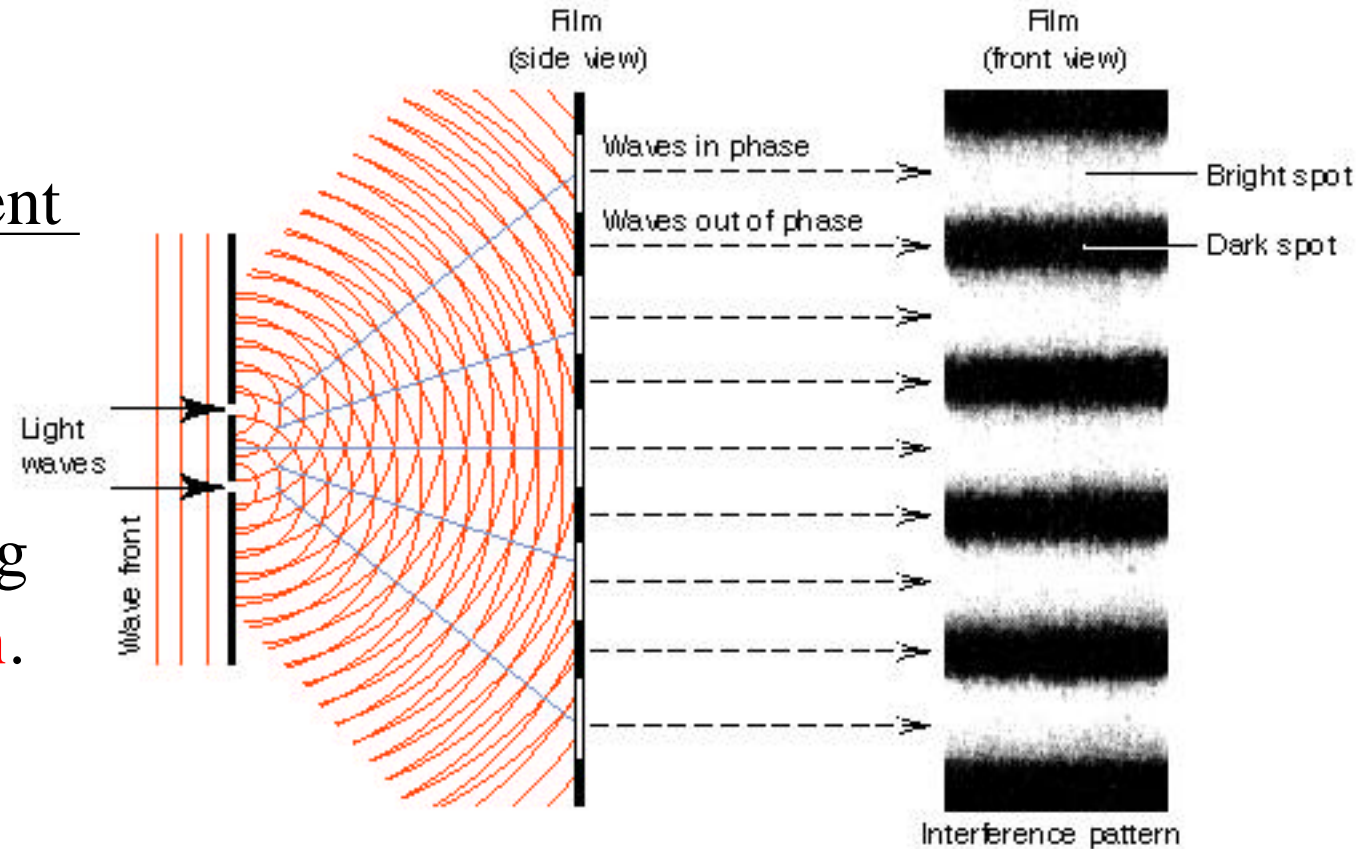
Concept 7-2. How particles and waves differ in terms of the phenomena of refraction, diffraction, and interference

Wave Diffraction

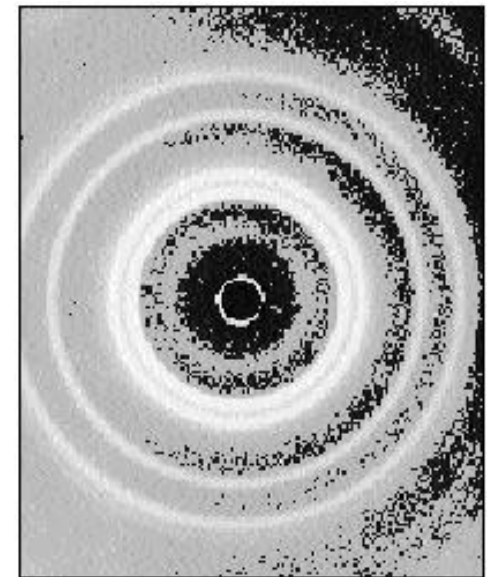
- If a wave passes through a slit about as wide as its wavelength, it forms a semi-circular wave on the other side of the opening.
- A stream of particles aimed at a small opening behaves quite differently:
- Some particles hit the edge and stop; those going through continue linearly in a narrower **stream**.



- If light waves pass through two adjacent slits, the emerging circular waves interact through **interference** creating a **diffraction pattern**.



- A diffraction pattern produced by X-rays.



The Particulate Nature of Light

Blackbody radiation

- As an object is heated, it begins to emit visible light, first as a red glow, then orange, then white light.
- Attempts to use electromagnetic theory to predict wavelengths of the emitted light failed.



- Max Planck proposed that the hot, glowing object could emit (or absorb) only certain amounts of energy: $E = nh\nu$

-where E is the radiation energy, ν is its frequency, n is a positive integer called a quantum number, and h is Planck's constant.

- With E in joules (J) and frequency in s^{-1} , h has units of $\text{J}\cdot\text{s}$:

$$h = 6.626 \times 10^{-34} \text{ J} \cdot \text{s}$$

Concept 7-3. The quantization of energy and the fact that an atom changes its energy by emitting or absorbing quanta of radiation

- If atoms can emit only fixed amounts of energy, it follows that atoms can have only **fixed electronic energy values**: $E = 1h, 2h, \dots$
- This constraint means that the energy of an atom is not continuous but **quantized**: it can exist only in certain fixed amounts.
- Each change in energy results from a “packet” of energy being gained or lost by the atom. This energy packet is called a **quantum**.
- Thus, an atom changes its energy state by emitting or absorbing a quantum of energy.

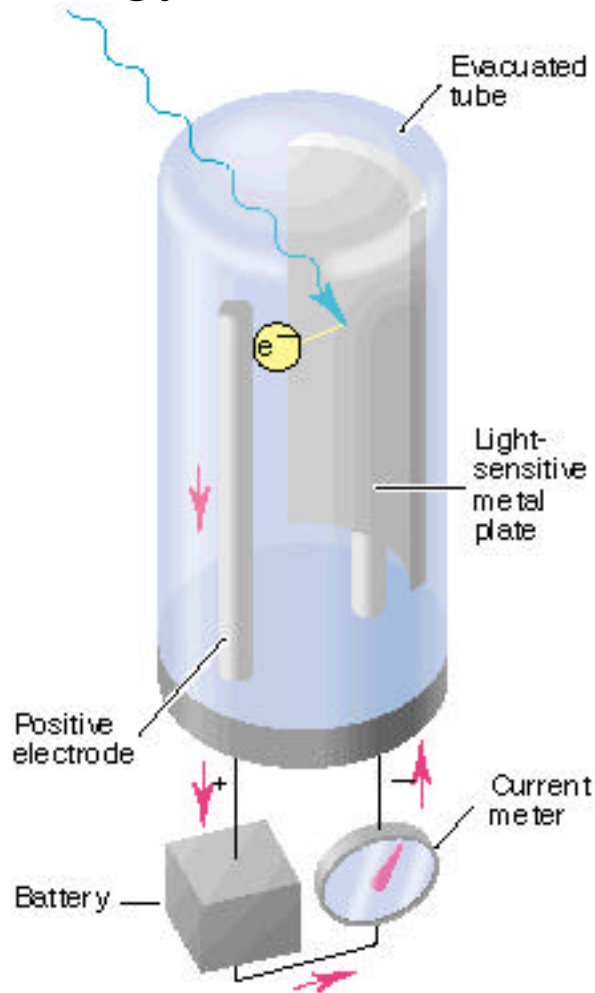
$$\Delta E_{\text{atom}} = E_{\text{emitted (or absorbed) radiation}} = \Delta nh\nu$$

- The smallest possible energy change for an atom is from one energy state to an adjacent one ($n = 1$):

$$\Delta E = h\nu$$

Concept 7-4. How the photon theory explains the photoelectric effect; the relation between photon absorbed and electron released

The photoelectric effect- when monochromatic light of sufficient energy hits the metal plate, an electric current flows.



1. Presence of a **threshold frequency**. Light shining on the metal must have a minimum frequency (which varies with the metal), or no current flows.
2. Absence of a time lag. Current flows the moment that light of high enough frequency shines on the metal, regardless of its intensity.

- Albert Einstein offered his **photon theory** to help explain the photoelectric effect.
- He proposed that radiation is **particulate**, occurring as quanta of electromagnetic energy, later called **photons**.
- An atom changes its energy when it absorbs/emits a photon, a “piece” of light whose energy is fixed by its frequency:

$$E_{\text{photon}} = h\nu = \Delta E_{\text{atom}}$$

How does Einstein’s photon theory explain the photoelectric effect?

1. Presence of a threshold frequency:

- A beam of light is composed of enormous numbers of photons.
- An electron is freed only when one photon of a certain minimum energy is absorbed. Energy depends on frequency ($h\nu$), so a threshold frequency is to be expected.

2. Absence of a time lag. An electron is freed the instant the atom absorbs a photon of enough energy. It doesn't accumulate energy.

- Current will be less in dim light than in bright light since fewer photons with enough energy eject fewer electrons per unit time.

Skill 7-2. Interconverting the energy of a photon with its frequency and/or wavelength

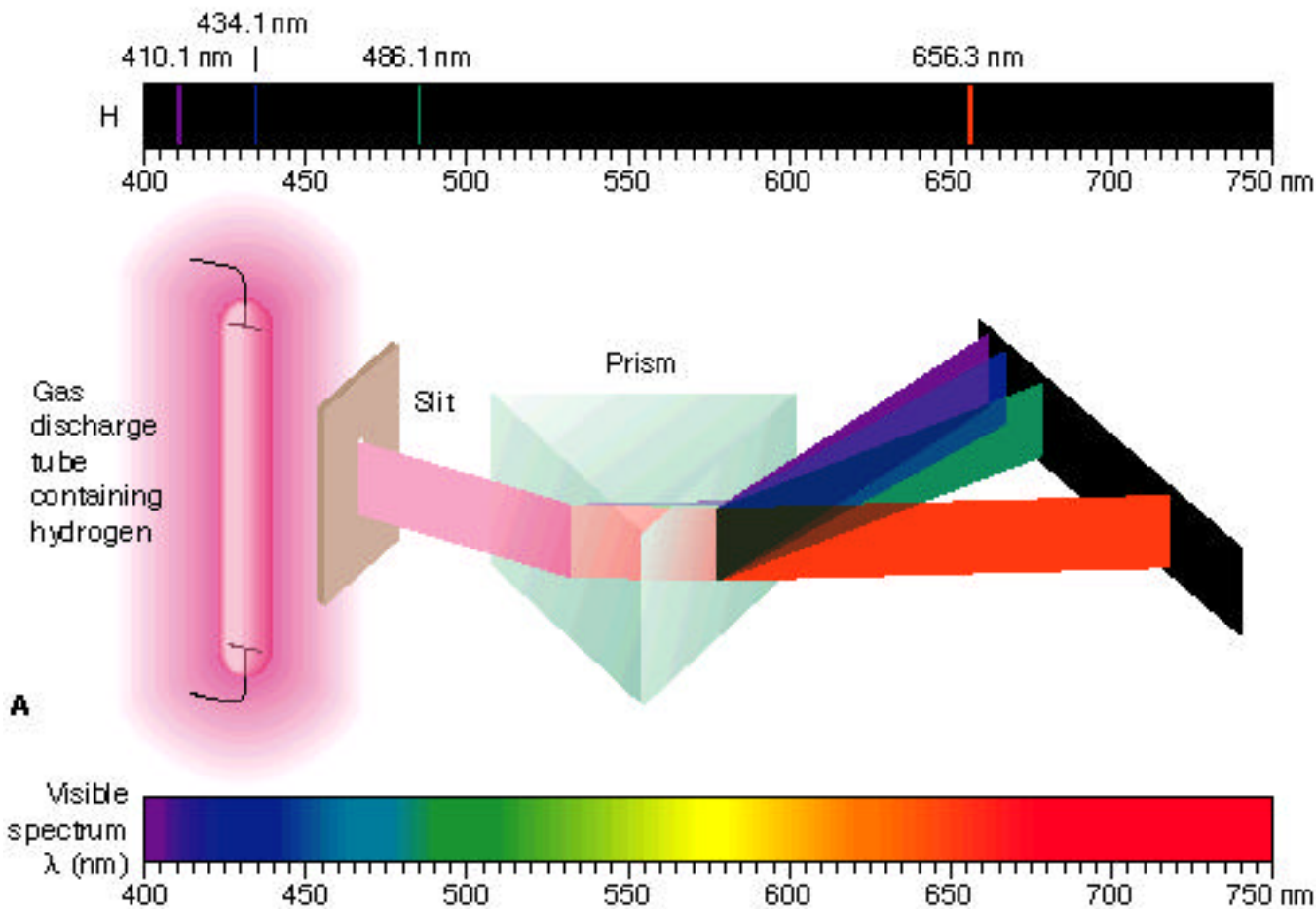
Problem: A microwave oven has a frequency of $2.45 \times 10^9 \text{ s}^{-1}$. What is the **energy** of one photon of this microwave radiation?

Solution: $E = h\nu = (6.626 \times 10^{-34} \text{ J} \cdot \text{s})(2.45 \times 10^9 \text{ s}^{-1})$
 $= 1.62 \times 10^{-24} \text{ J}$

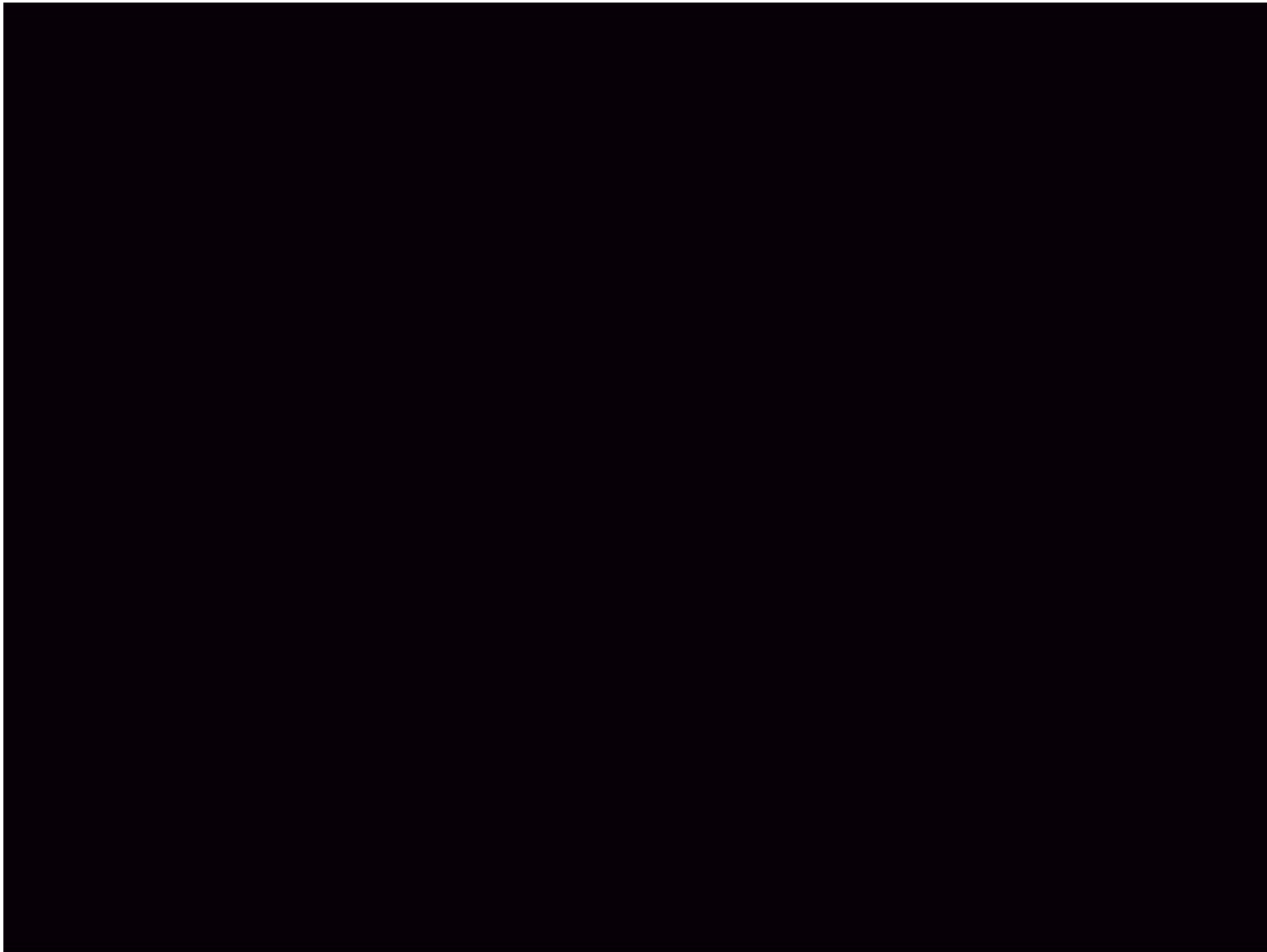
- The quantum and photon theories ascribed features to radiation that had been reserved for matter: **fixed amount** and **discrete particles**.

Concept 7-5. How the Bohr theory explained line spectra of the H atom; the importance of discrete atomic energy levels

- When an element is vaporized and thermally or electrically excited, it emits light. Hydrogen gas gives a pinkish color.
- Produces a **line spectrum**, a series of fine lines of individual colors.

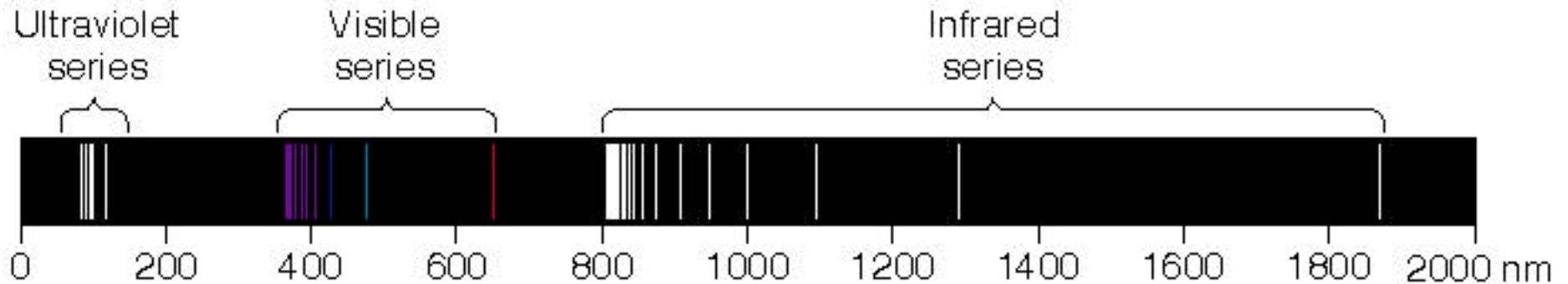
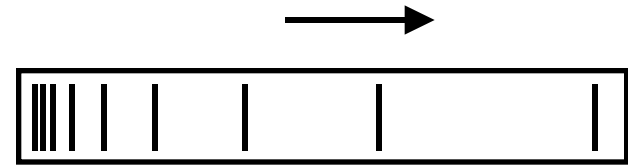


Atomic spectra video clip by Philip Morrison



- Spectroscopists discovered several series of H spectral lines in other regions of the electromagnetic spectrum.

- Each series showed a similar pattern.



- Equations of the form below, (Rydberg equation) were found to predict the position and wavelength of any line in a given series:

$$\frac{1}{\lambda} = R \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

Where n_1 and n_2 are positive integers with $n_2 > n_1$, and R is the Rydberg constant ($1.096776 \times 10^7 \text{ m}^{-1}$).

- For the series of spectral lines in the visible range, $n_1 = 2$:

$$\frac{1}{\lambda} = R \left(\frac{1}{2^2} - \frac{1}{n_2^2} \right), \text{ with } n_2 = 3, 4, 5, \dots$$

- Rydberg equation is **empirical**, based on data rather than theory.

Why do the spectral lines of hydrogen follow this pattern?

- Rutherford's nuclear model did not predict the existence of atomic line spectra. It had to be modified or replaced.



Niels Bohr

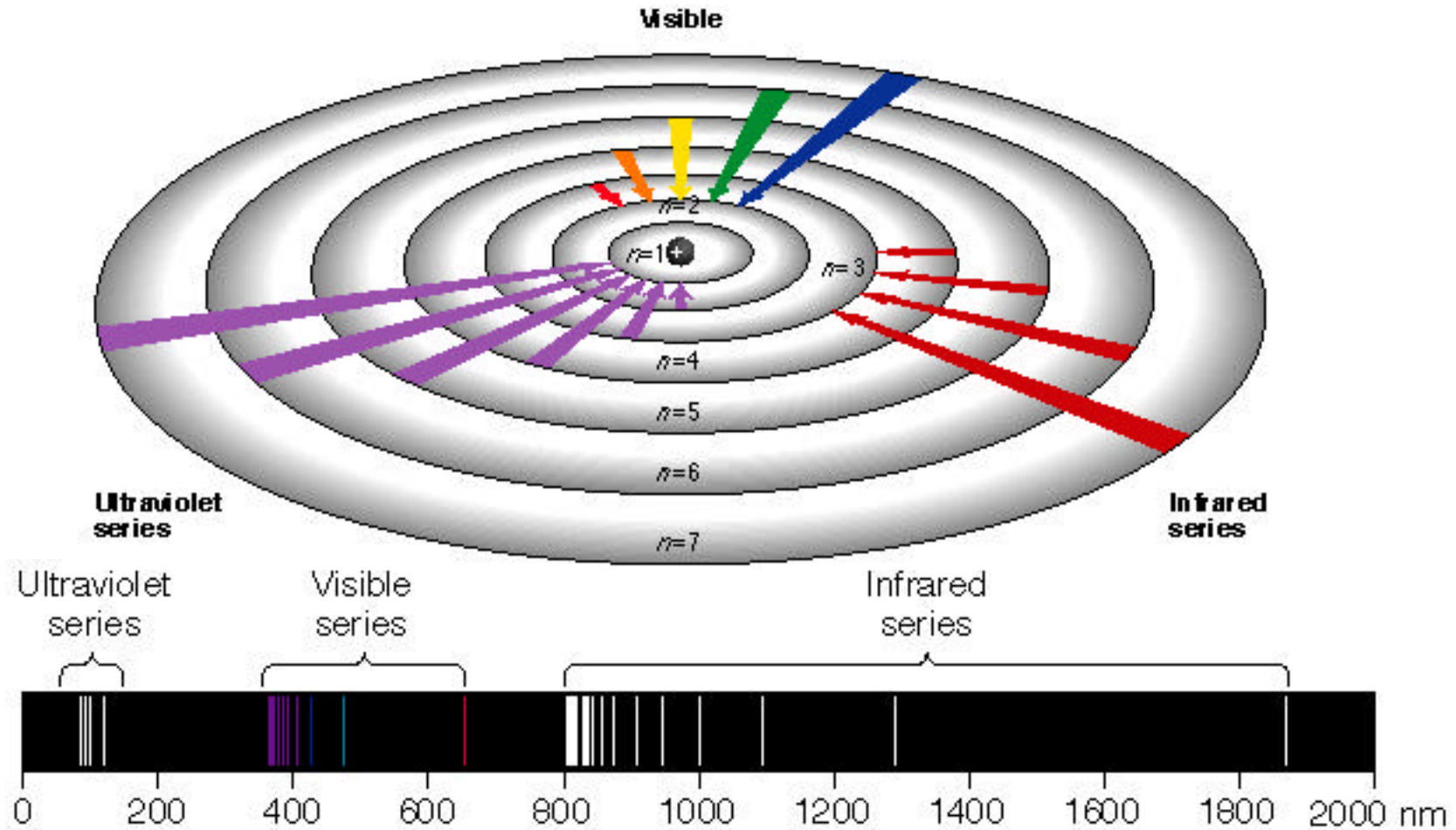
- Niels Bohr, working in Rutherford's laboratory, suggested a model for the hydrogen atom that predicted the observed line spectra.

The Bohr Model of the Hydrogen Atom

- Bohr used Planck's and Einstein's ideas about quantized energy proposing that hydrogen atoms had only certain fixed energy states.
- Each of these states was associated with a fixed circular orbit of the electron around the nucleus.
- Bohr proposed that atoms **do not radiate energy while in one of their fixed (stationary) energy states.**
- When the electron moves to a different orbit, the atom changes to another energy state.
- This happens only by the atom **absorbing** or **emitting** a photon.
- The photon energy equals the difference in energy between the two stationary states:

$$E_{\text{photon}} = \Delta E_{\text{stationary states}} = h\nu$$

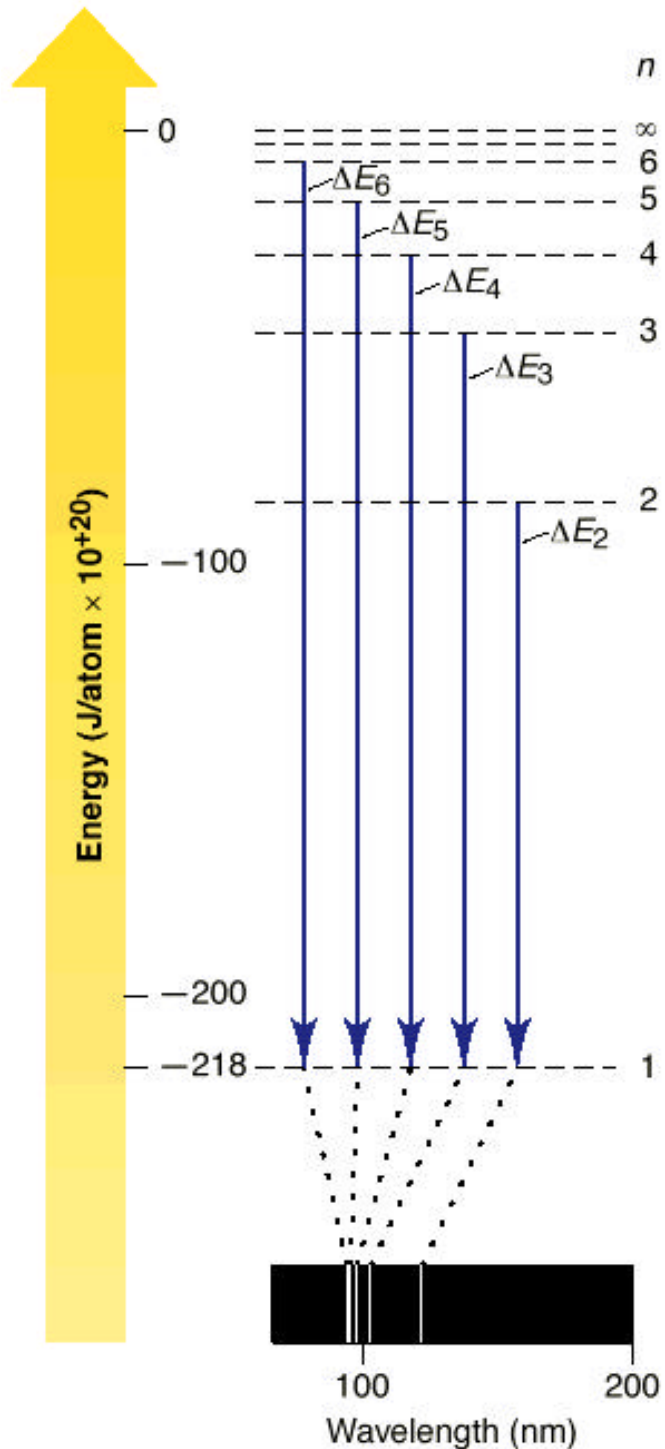
- Spectral lines represent **emission** of a photon of specific energy (and wavelength) as an electron **drops** from a higher energy state to a lower one.



- An atomic spectrum consists of lines instead of a continuum because the atom's energy has **only certain discrete levels**, or states.

- In Bohr's model, the quantum number n determines the **radius** of the electron's orbit and is directly related to the **atom's energy**.
- The lower the value of n , the smaller is the radius of the orbit and the lower is the energy level of the atom.
- When the electron is in the $n = 1$ orbit, the atom is in its lowest (first) energy level, called the **ground state**.
- The second energy level (second stationary state) and all higher levels are called **excited states**.
- A hydrogen atom in the 2nd energy level can return to the ground state by emitting a photon of a particular frequency:

$$E_{\text{photon}} = h\nu = E_{\text{first excited state}} - E_{\text{ground state}}$$



- When a sample of atomic hydrogen absorbs energy, different H atoms absorb different amounts.
- In various atoms all the allowable energy levels (orbits) are populated by electrons.
- Transitions from outer orbits to $n = 3$ gives the **infrared series** of spectral lines.
- Transitions from outer orbits to $n = 2$ gives the **visible series** of spectral lines.
- Transitions from outer orbits to $n = 1$ gives the **ultraviolet series** of lines.

The Energy States of the Hydrogen Atom

- Bohr obtained an equation giving the energies of the stationary states of the hydrogen atom:

$$E = -2.18 \times 10^{-18} \text{ J} \left(\frac{1}{n^2} \right)$$

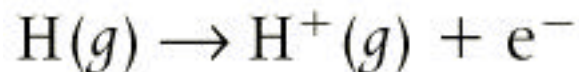
- The energy of the ground state ($n = 1$) is:

$$E = -2.18 \times 10^{-18} \text{ J} \left(\frac{1}{1^2} \right) = -2.18 \times 10^{-18} \text{ J}$$

- He defined zero energy ($E = 0$) as the state when the electron is totally removed from the nucleus; in other words, when $n = \infty$.
- So $E < 0$ (i.e. negative) for any smaller n values.
- The energy **difference** between any two energy levels is given by:

$$\Delta E = E_{\text{final}} - E_{\text{initial}} = -2.18 \times 10^{-18} \text{ J} \left(\frac{1}{n_{\text{final}}^2} - \frac{1}{n_{\text{initial}}^2} \right)$$

How much energy is needed to completely remove the electron from the hydrogen atom; that is, what is E for the following change?



Calculate this by substituting $n_{\text{final}} = \infty$ and $n_{\text{initial}} = 1$:

$$\begin{aligned}\Delta E &= E_{\text{final}} - E_{\text{initial}} = -2.18 \times 10^{-18} \text{ J} \left(\frac{1}{\infty^2} - \frac{1}{1^2} \right) \\ &= 2.18 \times 10^{-18} \text{ J}\end{aligned}$$

- E is positive because energy must be absorbed to remove the electron from the positive nucleus.

For one mole of H atoms,

$$\begin{aligned}\Delta E &= \left(2.18 \times 10^{-18} \frac{\text{J}}{\text{atom}} \right) \left(6.022 \times 10^{23} \frac{\text{atoms}}{\text{mol}} \right) \left(\frac{1 \text{ kJ}}{10^3 \text{ J}} \right) \\ &= 1.31 \times 10^3 \text{ kJ/mol}\end{aligned}$$

- This is the ionization energy of the hydrogen atom.

- Despite its success in accounting for spectral lines of the H atom, the Bohr model **failed** to predict the spectrum of any other element.
- The model worked well for one-electron species, but not for atoms or ions with more than one electron.
- The existence of discrete atomic energy levels is retained from Bohr's model in the current atomic model.

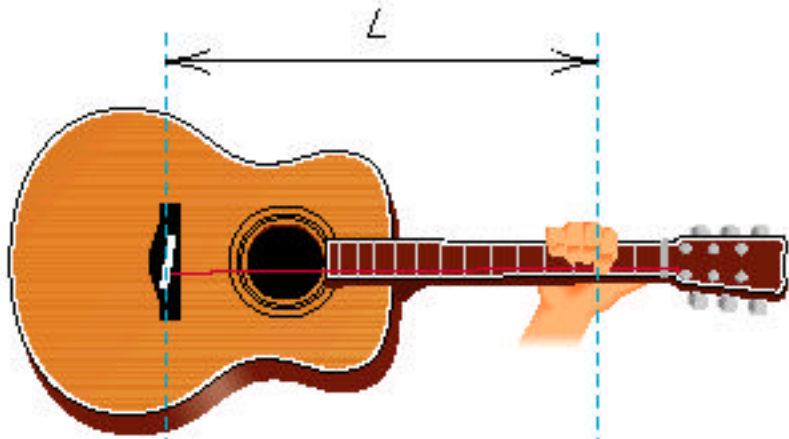
Concept 7-6. The wave-particle duality of matter and energy and the relevant theories and experiments that led to it



Louis de Broglie

The Wave-Particle Duality of Matter and Energy

- Louis de Broglie proposed a surprising reason for the existence of fixed energy levels in Bohr's model atoms.
- He said atoms are like other systems that display only certain allowed states, such as the wave created by a **plucked guitar string**.



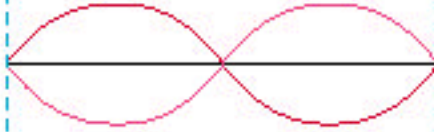
$n=1$



$L = 1\left(\frac{\lambda}{2}\right)$

1 half-wavelength

$n=2$



$L = 2\left(\frac{\lambda}{2}\right)$

2 half-wavelengths

$n=3$

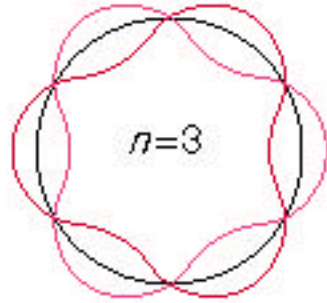


$L = 3\left(\frac{\lambda}{2}\right)$

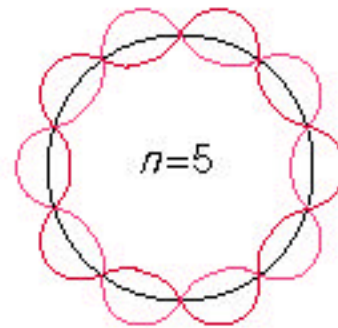
3 half-wavelengths

A

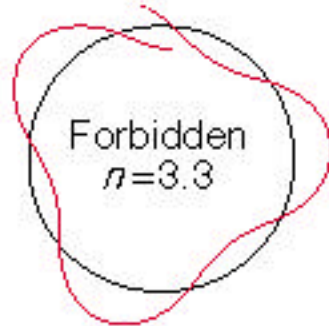
$L = n\left(\frac{\lambda}{2}\right)$



$n=3$



$n=5$



Forbidden
 $n=3.3$

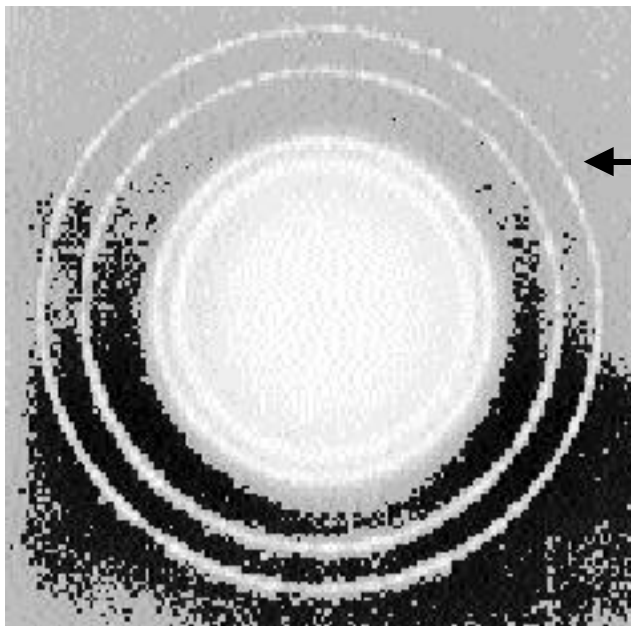
B

- DeBroglie suggested suggested that particles of matter such as electrons have **some properties of waves**.
- If electrons were wavelike and restricted to orbits of fixed radii, only certain possible frequencies and energies would be allowed.
- From mass-energy equivalence ($E = mc^2$) and the equation for the energy of a photon ($E = h\nu = hc/\lambda$), de Broglie derived:

$$\lambda = \frac{h}{mu} \quad \text{where } m \text{ is the mass and } u \text{ is the velocity.}$$

- An object's wavelength is inversely proportional to its mass; heavy objects have wavelengths that are very much smaller than the object
- Fast moving electrons have wavelengths close to the size of an atom ($\sim 10^{-10}$ m).

- If de Broglie's concept is correct, electrons should exhibit the wave properties of diffraction and interference.
- A good “slit” could be the spacings between atoms in a crystal.
- In 1927, C. Davisson and L. Germer guided a beam of electrons at a nickel crystal and obtained an **electron diffraction pattern**.



← Diffraction of an **electron beam** by thin aluminum foil.

- This observation proves that electrons behave as waves under some conditions.

Concept 7-7. The meaning of the uncertainty principle and how it limits our knowledge of electron properties

- If an electron has both particle and wave properties, its position in an atom can't be readily determined.



- Werner Heisenberg postulated the **uncertainty principle**: it is impossible to know simultaneously the exact position and velocity of a particle like an electron.

- The uncertainty principle is expressed mathematically as:

$$\Delta x \cdot m\Delta u \geq \frac{h}{4\pi}$$

x is the uncertainty in position.

u is the uncertainty in velocity.

- We cannot prescribe exact paths for electrons, such as the circular orbits of Bohr's model.
- At best, the **probability** of finding an electron in a **given volume** of space can be determined.

The Quantum-Mechanical Model of the Atom

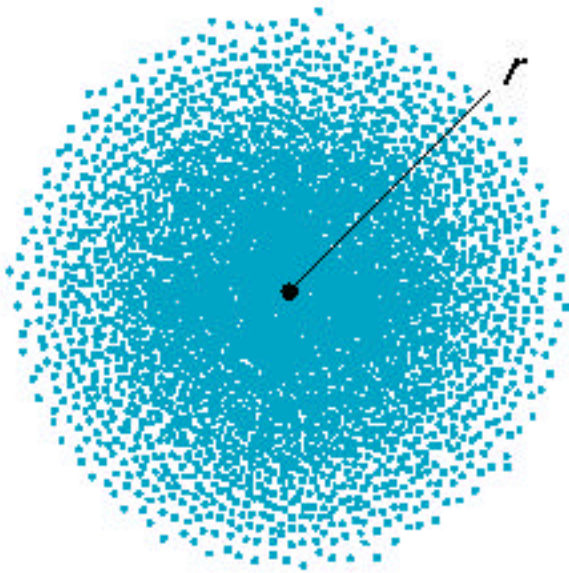


Erwin
Schrödinger

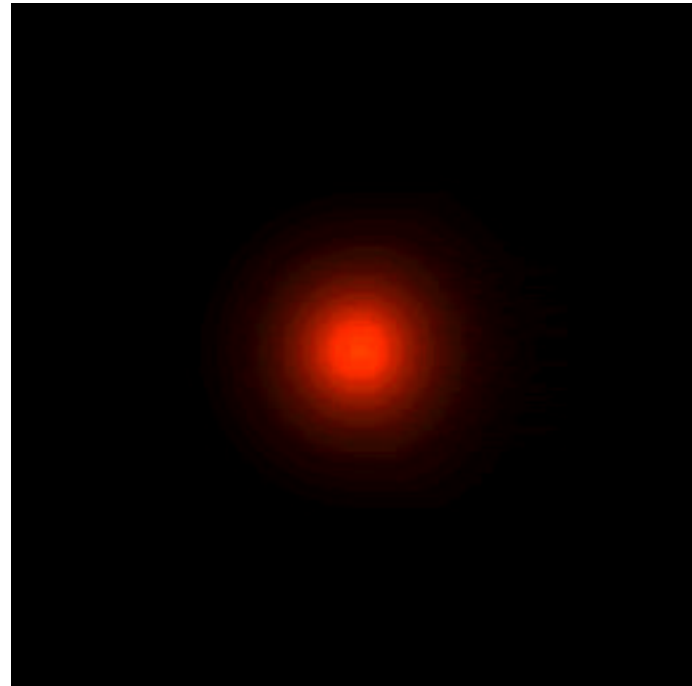
- Applying wave mathematics to the **electron wave**, Erwin Schrödinger derived an equation that is the basis for the quantum-mechanical model of hydrogen atom.
- The allowed wave-like motion of the electron leads to an atom with certain **fixed energy states** much like Bohr assumed.
- The electron's exact location cannot be determined.
- Solutions of Schrödinger's wave equation are functions, ψ , that describe **atomic orbitals**.
- Each ψ describes a fixed-energy state the electron can occupy and ψ^2 gives the probability of finding it in a given 3-dimensional space

Concept 7-8. The distinction between ψ (atomic orbital) and ψ^2 (probability of the electron being at a distance r from the nucleus)

- This probability can be shown pictorially by means of an electron probability density diagram, or simply, an electron density diagram.
- Electron probability density in the H atom ground state.

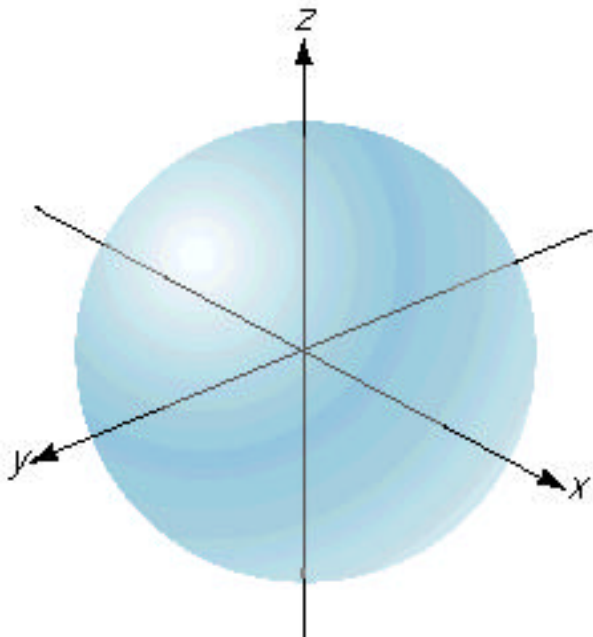


electron “cloud”

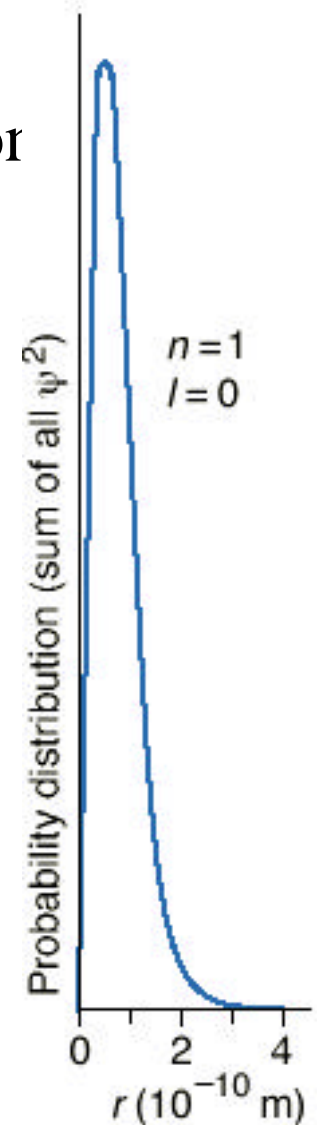


Concept 7-9. How electron density diagrams and radial distribution plots depict the probability of electron location within the atom

- Dividing the volume of atom into thin, concentric, spherical layers and calculating ψ^2 in each, the variation of electron density with **radial distance** is found.
- The probability of the electron being far from the nucleus is very small, but not zero.



A 90% contour depiction of the ground state; the volume in which the electron spends 90% of its time.



- The maximum radial probability for the ground-state of H atom appears at (0.529 Å, or 5.29×10^{-10} m), the same as 1st Bohr orbit.
- The electron spends most of its time at the same distance that the Bohr model predicted it spent all of its time.
- Each atomic **orbital**, ψ , has a distinctive radial probability distribution and probability contour diagram.

Concept 7-10. The hierarchy of the quantum numbers that describe the size (n , energy), shape (l), and orientation (m_l) of an orbital

Quantum Numbers of an Atomic Orbital

- An atomic orbital is specified by **three quantum numbers**.
- These determine the orbital's **size, shape, and orientation** in space.

Concept 7-11. The distinction between level (shell), sublevel (subshell), and orbital

- The principal quantum number (n) is a positive integer (1, 2, 3, etc).

n indicates the relative **size** of the orbital (the distance from the nucleus of the peak in the radial probability distribution plot).

n specifies the **energy level** of the H atom: the higher the n value the greater the energy.

- The azimuthal quantum number, l , is an integer from **0 to $n - 1$** .

l is related to the **shape** of the orbital.

For orbitals with $n = 2$, l can have a value of 0 or 1; with $n = 3$, l can be 0, 1, or 2; etc.

The number of possible l values equals the value of n .

- The magnetic quantum number (m_l) is an integer from $-l$ through **0** to $+l$.

m_l prescribes the **orientation** of the orbital in the three-dimensional space about the nucleus.

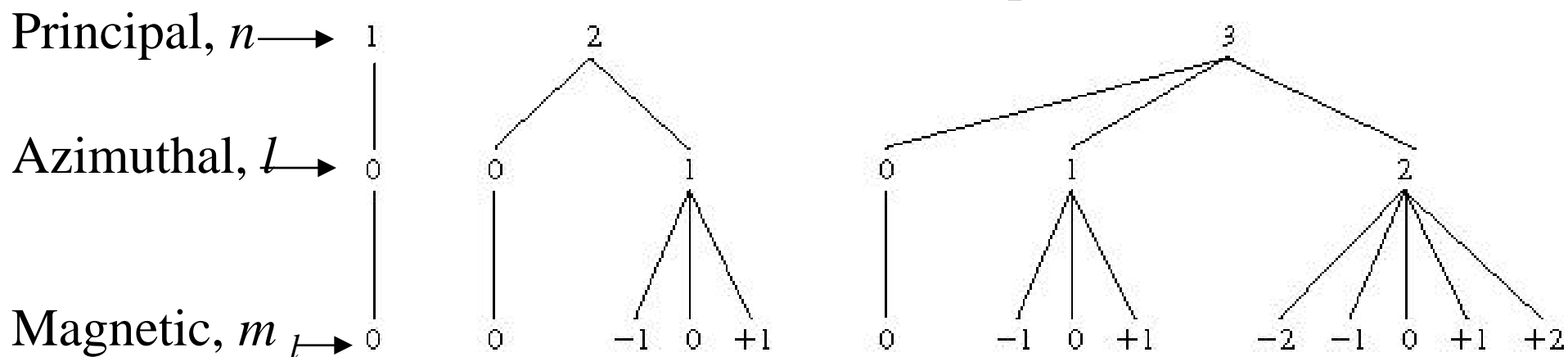
The possible values of an orbital's magnetic quantum number are set by its azimuthal quantum number.

An orbital with $l = 0$ can have only $m_l = 0$.

An orbital with $l = 1$ can have m_l values of -1 , 0 , or $+1$.

The number of possible m_l values, or orbitals, for a given l value is **$2l + 1$** .

Allowed values of quantum numbers



Determining Quantum Numbers for a Given Energy Level

Problem: (a) What values of the azimuthal and magnetic quantum numbers are allowed for a principal quantum number of 3?

(b) How many orbitals are allowed for $n = 3$?

Solution: Determining l values:

l takes values from 0 to $n-1$. For $n = 3$, $l = 0, 1, 2$.

Determining m_l for each l value:

For $l = 0$, $m_l = 0$

For $l = 1$, $m_l = +1, 0, -1$

For $l = 2$, $m_l = +2, +1, 0, -1, -2$

Since there are **nine** m_l values, there are **nine orbitals with $n = 3$** .

- The total number of orbitals for a given n value is n^2 .
- The atom's energy levels, or **shells**, are given by the n value.
- The $n = 1$ shell has lower energy and a greater probability of the electron being closer to the nucleus than higher shells.
- The atom's sublevels, or subshells, are given by the n and l values. Each shell contains **subshells** that designate the shape of the orbital.

- Each subshell is designated by a letter:
 - $l = 0$ indicates an **s subshell**.
 - $l = 1$ indicates a **p subshell**.
 - $l = 2$ indicates a **d subshell**.
 - $l = 3$ indicates an **f subshell**.
- Subshells are named by using the n value and the letter designation.

The subshell with $n = 2$ and $l = 0$ is called the **2s** subshell.

The only orbital in the 2s subshell has $n = 2$, $l = 0$, and $m_l = 0$.

The subshell with $n = 3$ and $l = 1$ is the **3p** subshell.

It has three orbitals: $n = 3$, $l = 1$, and $m_l = +1, 0$ and -1 .
- Each of an atom's orbitals is specified by a set of **n , l , and m_l** values
- These fix the orbital's **size, energy, shape, and spatial orientation**.

Skill 7-5. Determining quantum numbers and sublevel designations

Problem: Give the name, magnetic quantum numbers, and number of orbitals for each subshell with the quantum numbers shown:

(a) $n = 3, l = 2$ (b) $n = 2, l = 0$ (c) $n = 5, l = 1$ (d) $n = 4, l = 3$

Solution:

	<u>n</u>	<u>l</u>	<u>subshell name</u>	<u>possible m_l</u>	<u>No. of orb.</u>
(a)	3	2	3d	-2, -1, 0, +1, +2	5
(b)	2	0	2s	0	1
(c)	5	1	5p	-1, 0, +1	3
(d)	4	3	4f	-3, -2, -1, 0, +1, +2, +3	7

Identifying Incorrect Quantum Numbers

Problem: What is wrong with each of the following quantum number designations and/or subshell names?

	<u>n</u>	<u>l</u>	<u>m_l</u>	<u>NAME</u>
(a)	1	1	0	1 <i>p</i>
(b)	4	3	+1	4 <i>d</i>
(c)	3	1	-2	3 <i>p</i>

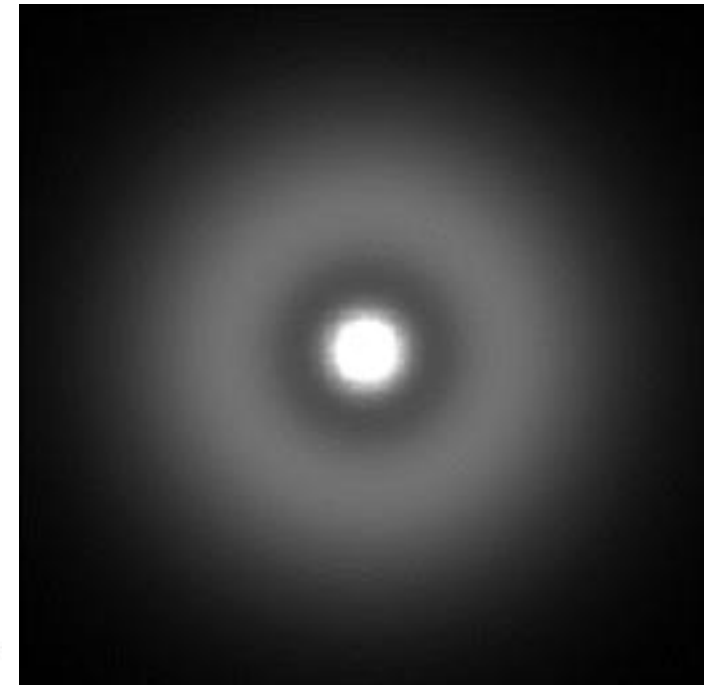
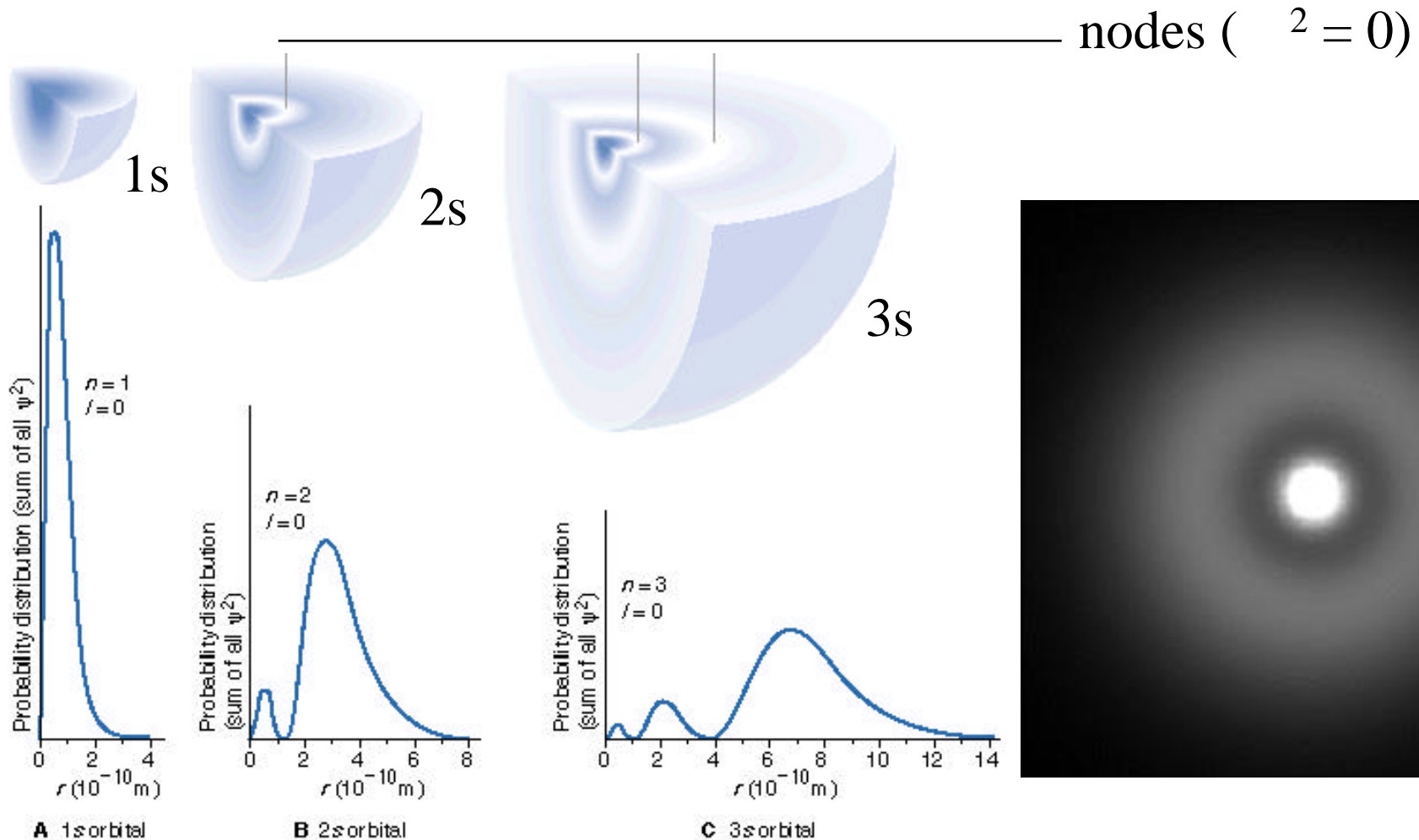
Solution: (a) A sublevel of $n = 1$ can have only $l = 0$, not $l = 1$. The only possible subshell is 1s.

(b) A subshell with $l = 3$ is an f subshell, not a d subshell. The subshell name should be 4f.

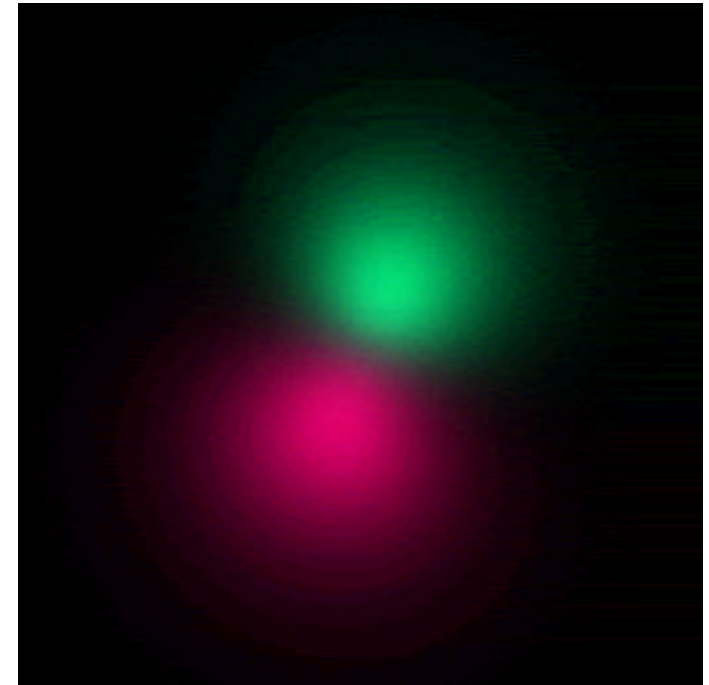
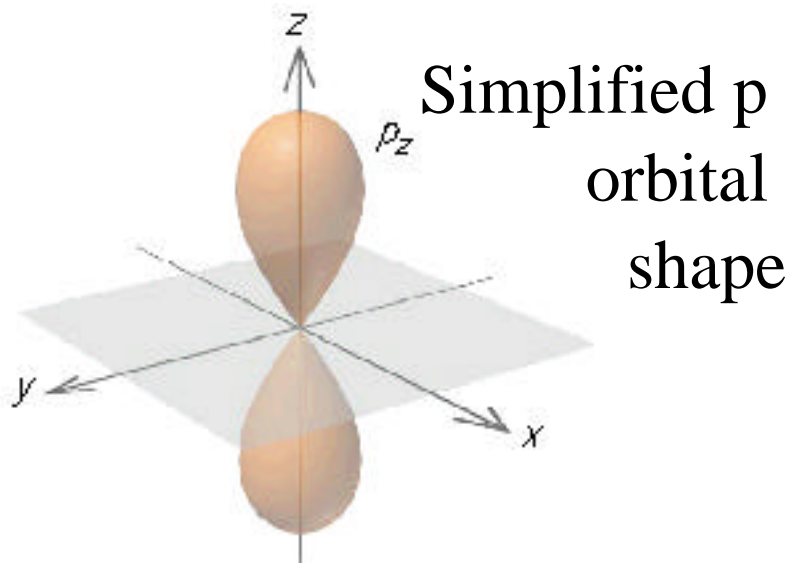
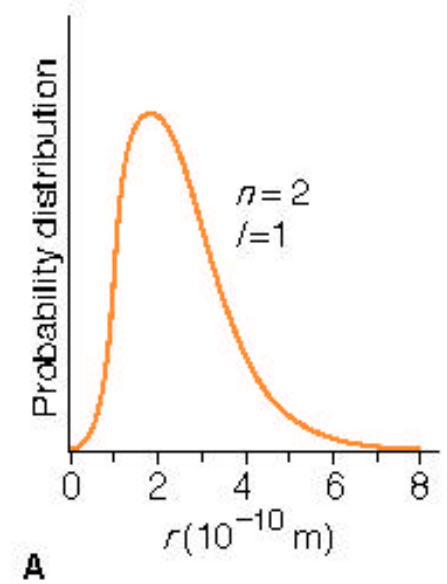
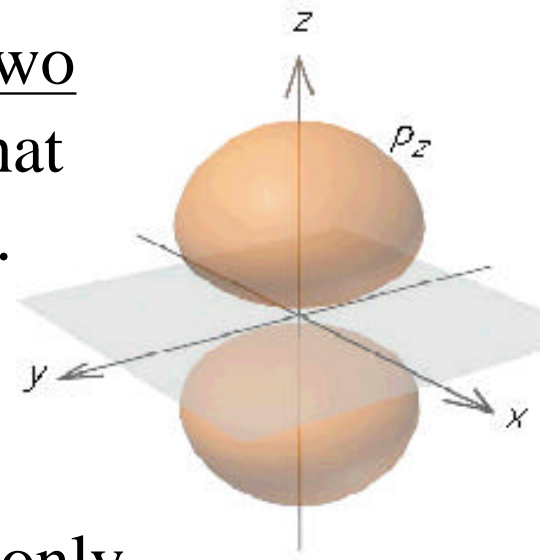
(c) A subshell with $l = 1$ can have only m_l of -1, 0, +1, not -2.

Concept 7-12. The shapes and nodes of s, p, and d orbitals

- Each subshell of the H atom has orbitals with a characteristic shape
s orbitals ($l = 0$) are spherically symmetrical around the nucleus.

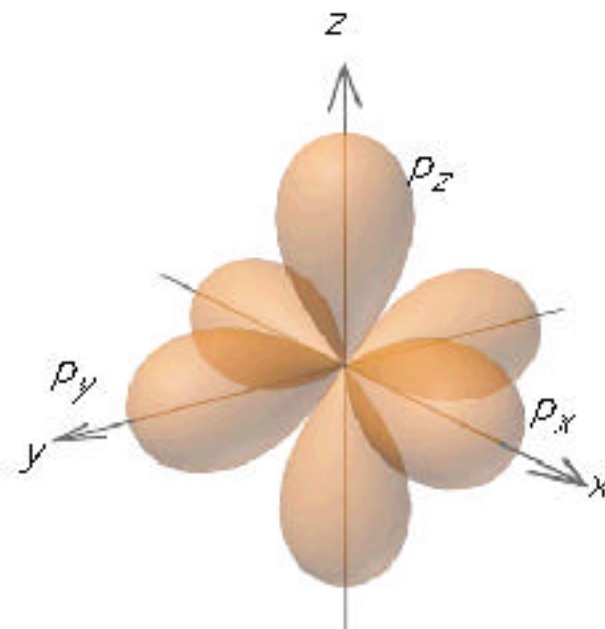


- The p orbital (with $l = 1$) has two regions of higher probability that lie on either side of the nucleus.
- The nucleus lies at a **nodal plane** of p orbitals.
- The max. value of l is $n - 1$, so only shells with $n \geq 2$ can have p orbitals.
- Energies of p orbitals: $2p < 3p < 4p \dots$



- The three possible m_l values: +1, 0, and -1, which lead to three mutually perpendicular p orbitals. p_x p_y p_z

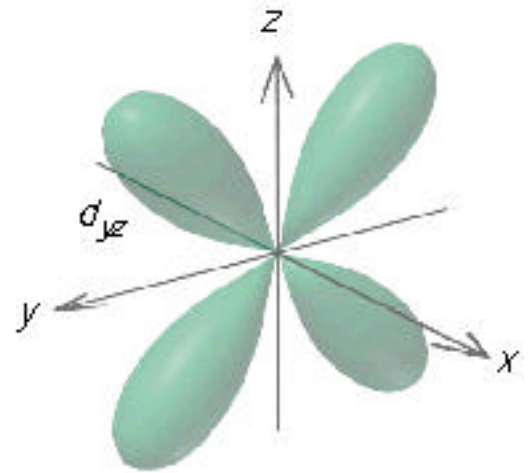
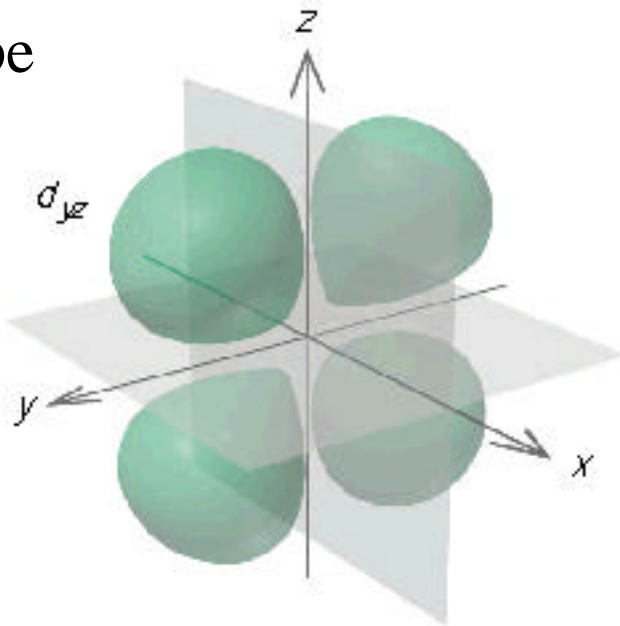
- The three p orbitals are equivalent in size, shape, and energy, differing only in orientation.



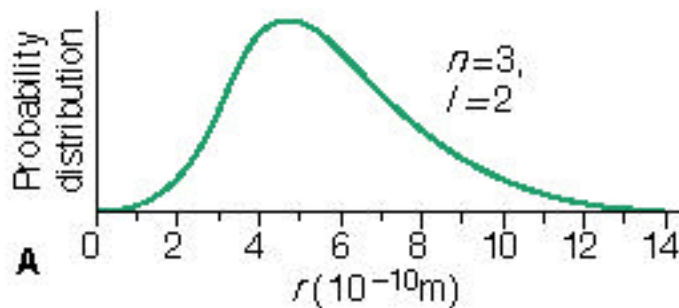
The d orbitals

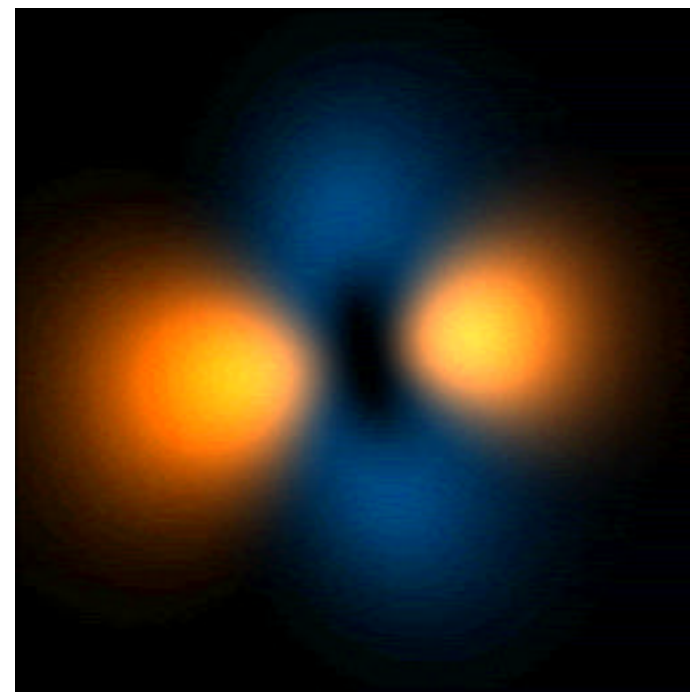
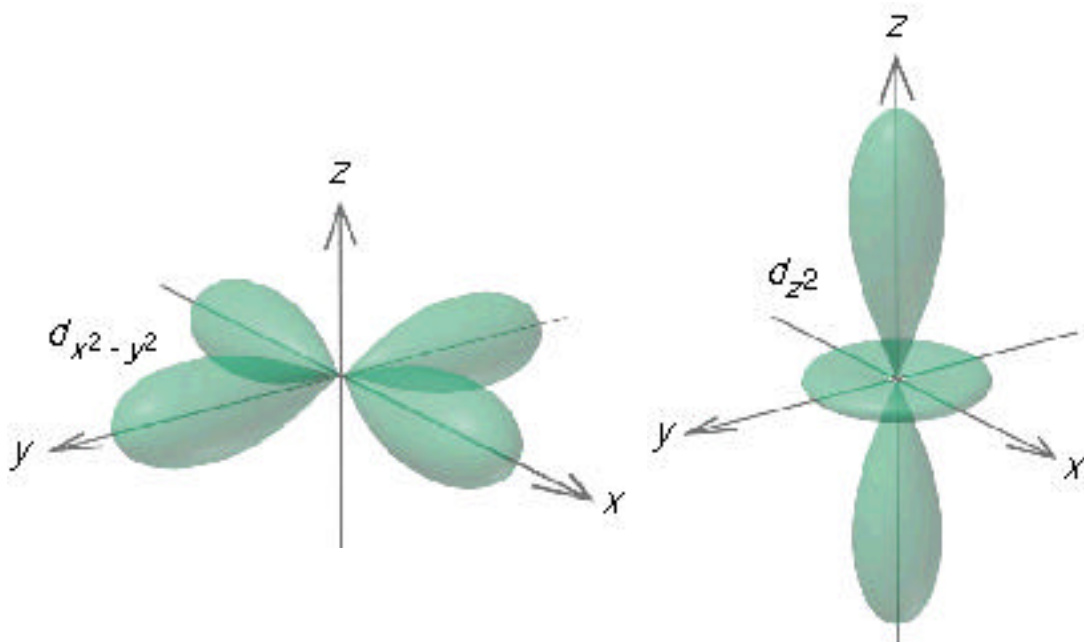
- Orbitals with $l = 2$ are called d orbitals. There are **five** possible m_l values for the $l = 2$ sublevel: +2, +1, 0, -1, and -2.
- d orbitals can have any one of **five** different orientations.

Actual shape
of d_{yz}



simplified
shape of d_{yz}





- In accord with the quantum number rules, a d orbital must have a principal quantum number of at least $n = 3$.
- The 4d orbitals extend farther from the nucleus than the 3d, and the 5d orbitals extend still farther.

Energy Levels of the Hydrogen Atom

